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Meteorology**3711 Chemical composition and chemical interactions: RECENT MEASUREMENTS OF THE DISTRIBUTION OF VARIOUS VAPOR IN THE STRATOSPHERE AT HIGH LATITUDES**

E. O'Brien and V.A.J. Evans, Atmospheric Research Division, 1990 University Street,

Downsville, Ontario, N3L 3T6

The measurements of the altitude distribution of stratospheric water vapor have been made with a long-distance microwave photometer which measures long-wavelength absorption in the solar beam at the 2.1 μ m spectral region.

The photometer was flown from Cape Parry, Canada, at 69°N, 100°W (latitude 70.4°N, on

December 4, 1979) and from Fairbanks, Alaska (latitude 62.9°N, longitude 151.7°W), on March

13, 1980, at sunset. These measurements indicate that the total water vapor mixing ratio increases from approximately 1.5 ppm at 50 km to approximately 8 ppm at 90 km and then decreases at higher altitudes.

These features are consistent with theoretical considerations of the stratospheric water vapor distribution and with other experimental investigations. A source region from which the water vapor at 50 km altitude and east over the tropics are expected. A comparison of laboratory measurements indicates that the step of the water vapor line in the 2.1 μ m spectral region is significant. This is important for the application of the water vapor line to the analysis of the atmospheric measurements.

J. Geophys. Res., Green, Paper 1C0926

3712 Chemical composition and chemical interactions:**THE PM OF MARINATE PRECIPITATION: A PRELIMINARY REPORT**

J. M. Miller (NOAA Air Resources Laboratories, Silver Spring, Maryland 20910)

A. M. Yudin (NOAA-National Sea Observatory, Hills, Hawaii 96746)

Marine precipitation samples have been collected at the National Weather Service at Maui since 1978. The collection altitude often ranged from sea level to 300 m. Samples were analyzed on a day of collection for pH and major ions. Dissolved particulate matter, such as sulfate and nitrate, were also selected samples during the period.

The data show progressive increases of acidity with altitude. The pH values averaged pH 5.1, in contrast to the values above 2500 m, which average pH 4.5. It is postulated that the acidity in the stratosphere might be due to the formation of sulfuric acid either natural or man-made, being transported long distances in the mid-stratosphere and being converted to the form of the resulting acids. (Precipitation chemistry; acid rain; long-range transport.)

J. Geophys. Res., Lau, Paper 1C0927

3713 Chemical composition and chemical interactions:**THE ANNUAL VARIATION OF STRATOSPHERIC CO₂ CONCENTRATION OBSERVED IN THE NORTHERN HEMISPHERE**

Y.-O. Boxall (Division of Atmospheric Physics, 902, Box 771, Macmillan, Victoria, Australia 3900)

P. J. Hines (Department of Meteorology, University of Melbourne, Victoria, Australia 3010)

Details of the annual variation of the stratospheric concentration of CO₂ are presented by Boxall, Hines, and Marshall (1981). The seasonal changes in the amplitude of the annual oscillation appear to have increased in recent years with the introduction of observations based on the mass spectrometer data, rather than the CO₂ data. This change is discussed in terms of changes in the atmospheric reprocessing and photochemistry and the use of CO₂ data. The analysis does not allow

for the separation of several possible causes of amplitude change. However, if the change is interpreted as reflecting enhanced atmospheric CO₂ uptake equivalent to an increase in the net uptake of carbon dioxide between 1970 and 1979, and a growth of the northern hemisphere stratospheric CO₂ of 0.5×10^{12} kg of carbon per year, the calculated CO₂ content with recent inventory studies which indicate that some CO₂ sources have acted as a net sink of 10^{12} kg of carbon per year in recent decades, (CO₂, annual cycle, seasonal change, carbon cycle).

J. Geophys. Res., Green, Paper 1C0928

3714 Chemical composition and chemical interactions:**SIMULATION OF NITROGEN CONSTITUENT MEASUREMENTS FROM THE 1976 STRATOSCOPE III FLIGHT**

V.F.J. Evans, C.T. McElroy, J.B. Kerr, Atmospheric Environment Service, 4905 Dufferin Street, Downsview, Ontario, M3J 2M5

J.C. McConnell, GRS, York University, 4700 Keele Street, Downsview, Ontario, M3J 2P3

A simulation of the altitude distribution of the nitrogen constituents measured on the STRATOSCOPE III flight is presented. The simulation was carried out using a fixed dependent stratospheric model with extensive photochemistry.

Temperature and temperature profiles were measured by the stratoscope. The measured ratios for NO/NO_2 and NO_2/NO are significantly larger than the measured ratios for NO/NO_2 and NO_2/NO in the altitude range 20 to 30 km. Since there is no direct way to relate the hydroxyl densities, it is proposed that the actual hydroxyl densities are a factor of 3 times those calculated in the model using current chemical reaction rates. The calculated pressure dependence of the NO/NO_2 reaction rate as well as including the NO_2/NO reaction rate in the model, reduce the NO densities in the model over this altitude range. An empirical simulation of the measured ratios is obtained. The pressure dependence chosen for the rates was that reasonable agreement with existing measurements of hydroxyl densities at higher altitudes was still obtained.

J. Geophys. Res., Green, Paper 1C0929

3715 Electrical phenomena:**BALLOON-BORNE MEASUREMENTS OF ELECTRICAL CONDUCTIVITY, MOBILITY AND THE RECOMBINATION COEFFICIENT**

J. M. Rosen (Dept. of Physics and Astronomy, University of Wyoming, Laramie, WY 82071) and D. R. Johnson

Balloon-borne measurements of the positive polar conductivity (σ_p) from near ground level to 35 km are presented. The data are obtained in two separate experiments. The measured conductivities of the ionization rate (λ) and the positive ion density (n_p). These three parameters are used to calculate the average positive ion mobility (μ_p) and the recombination coefficient (κ). Within the experimental accuracy ($\pm 10\%$) the reduced mobility is constant with altitude at a value of about $1.1 \times 10^{-2} \text{ cm}^2/\text{V s}$. The reduced recombination coefficient is found to be constant with altitude. Theoretical predictions favorably compare with the theoretical predictions but there is some serious question as to whether the theoretical values can in fact be compared directly to the experimental results obtained here.

J. Geophys. Res., Green, Paper 1C0930

3716 Electrical phenomena:**MEASUREMENT OF LOWER ATMOSPHERIC VERTICAL POTENTIAL DIFFERENCES**

R. H. Holwede (Space Sciences Laboratory, The Aerospace Corporation, P.O. Box 92911, Los Angeles, CA 90048), M. H. Dasey, E. K. Schubert, and C. Youngblut

A high impedance system has been developed to make direct measurements of the atmospheric electric field at altitudes of several thousand feet. A tethered balloon flown from Walllops Island, Virginia is used to lift a high voltage insulated wire and a conducting center electrode. A high voltage source is used to charge the electrodes. The balloon is released in October 1980. The balloon was equipped with a payload to measure wind speed and direction, and electric potential. The highest vertical potential difference was measured at 170,000 feet, or 1000 feet.

The short circuit currents which could be seen through the wire were in the 10 milliampere range and the impedance of the wire was measured to be about 10 ohms. This paper will describe the apparatus and details of these measurements. (Atmospheric potentials, tethered balloon).

J. Geophys. Res., Paper 1C0721

3717 Climatology:**ZONAL-WAVE PATTERNS IN THE STRATOSPHERE EAST OF THE NORTHERN HEMISPHERE DURING THE LAST TWO SUNSPOT CYCLES**

S. Neukirch (Institut für Meteorologie, Federal University Berlin, 1000 Berlin 33, FRG)

Stratospheric mean-height and temperature data of three levels are used.

Four-year time series for a 10- to 10-degree latitude band in the stratosphere are obtained, whose amplitudes

as well as the mean latitude and longitude of differences between successive years are investigated in the Carrington cycle during the same period.

An atmospheric electrical parameter has been

investigated in a high set of ΔT -heights.Production estimates at about $2300 \text{ K} \text{V} \text{ m}^{-1} \text{ sec}^{-1} \text{ rad}^{-1} \text{ cm}^{-2}$.

This leads to ion densities found in the lower stratosphere.

Electrical conductivity in the lower stratosphere is about 100 times above the mesosphere.

This result is attributed to a condensation effect.

This has been done for the first time.

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Copyright 1981 William Bowles Medalist Harriet Friedman in front of a rubbing of the Soviet Medal (see page 574).

Considerations in the Development of a National Geophysical Data Policy

Juan G. Roederer

Geophysical Institute
University of Alaska, Fairbanks**Introduction**

Science emerged when it became apparent that the images of the world and of environmental events, acquired through the senses and registered by the human brain in natural day-to-day experience, contained inaccuracies and subjective biases that interfered with the development of an increasingly complex society. Recognition of the need for systematic statistical verification of predictions and for unbiased reporting and recording of both successes and failures of predictions became the fundamental driving force in the development of the scientific method and scientific thought. It became apparent that in order to establish a repertoire of reliable information on cause-and-effect relationships, environmental exploration and documentation would have to be expanded from subjectively "relevant" phenomena to others that bore no direct relation to, or had no effect on, the human organism. It was also realized that a merely passive, qualitative, random observation of environmental events did not yield sufficient information. Active, quantitative probing and systematically planned experimentation became a necessity; the empirical method was born. Our sensory systems needed extension to achieve higher accuracy in the acquisition of environmental information, and scientific instruments were developed to make the measurements required for a quantitative description of environmental events over a wide range of domains. Finally, it was realized that the use of ex-ontological documentation, data, and information systems (books, data repositories, computers, etc.) was essential for organizing experimental paradigms, for their statistical interpretation, for recording results, and, in general, for the development of an "objective truth" about environmental events.

Since the end of World War II, human society in most advanced countries has undergone a profound transition from an "industrial society" to an "information society," in which industrial, economic, and military power is conditioned to information-processing power, and societal well-being, social organization, and government are conditioned to the information transfer capacity among elements of the political system. Bell (1973) and many others have described this transition.

Research and development in an information society is heavily problem oriented, with the basic ethic of "solving so-

cial's problems." A recent Arthur D. Little Inc. (1978) report described this period as the "problem-oriented Era III." Initially, in the discipline-oriented Era I, basic research and discipline-centered R&D were the main sources of new knowledge. Era I, however, persists, i.e., must persist, into Era III, for it provides the building stones on which problem-oriented information systems are to be built. The transition period, during the 1950's through the mid-1960's, has been called the "mission-oriented Era II," the basic ethic of which was to "organize to do a job," with data-intensive research efforts. Mission-oriented endeavors must also subside into Era III.

Era I information systems mainly handle "end products" of research (such as, articles, books, etc.); producers and users of information normally belong to the same discipline, and producers and users of raw data often belong to the same research group. In contrast, Era III information systems will mainly handle cross-disciplinary data flow (often intensive raw data flow) and deal with cooperative data analysis, which will become fundamental tools in the search for answers to the problems posed. Data producers and users usually belong to different groups and even to different disciplines, but they must be able to communicate with each other and work together in data analysis. The data needed are often of synoptic type, acquired in large monitoring networks, observatories, laboratory facilities, or based on large-scale statistics or surveys that cannot be operated or conducted by individual groups or institutes.

Today, research and development is becoming increasingly dependent on the availability of huge amounts of data and information stored in public repositories (data centers, technical libraries) accessible to users other than the originators. Yet the organization of data flow into and out of repositories has so far followed mostly "local" or disciplinary Era I or II needs, evolving as these needs arise, with little national and interdisciplinary coordination. Even within the disciplines, there exist only a few formal agreements, on the part of certain data originators, regarding formats, units, and type of data to be stored. In summary, a full transition to Era III information systems has not yet taken place.

The Arthur D. Little report identifies nine problem categories as fundamental targets in Era III: environment, energy, economic well-being, safety, public health, transportation, crime prevention and administration of justice, housing, and welfare. Three broad academic disciplines must cooperate in this endeavor, namely earth sciences, life sciences, and social sciences. Each one must develop Era III data and information systems which ultimately will provide, collectively, the quantitative answers sought for the problem categories above. Geophysics is now ready for the establishment of Era III-type data repositories. Indeed, large international programs such as the International Geophysical Year have contributed to a "data explosion," which has continued gathering momentum exponentially ever since. Hence geophysics should serve as a convenient proving ground for the testing and establishment of a national data policy appropriately tuned to the needs of tomorrow's society.

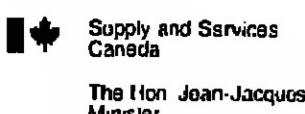
Data Infrastructure and Management Issues

Data storage—whatever the level of the data stored—necessitates the concomitant storage of information on the underlying data infrastructure, that is, information on the original physical magnitudes(s) measured, on the circumstances of the measurement, on format and units, and on the software, assumptions, models, etc., used in the data-processing stages. Without such information, the stored data is worthless. For level I data, such information is usually available only to the experimenter. Once this information is lost, raw data become worthless. For level II data, much less concurrent information is necessary. Often it is enough to know what physical magnitude they pertain, what the units are, and in what way the data varies over time (e.g., as a function of time, as a function of position, etc.). Level II data can be used by "secondaries" users, provided, however, that there is confidence in the data originators. If for some reason that confidence or credibility is lost, level II data become worthless. Essentially the same applies to level III data.

There are branches of physics in which the secondary user (a user that does not belong to the group that acquires the raw data) only needs level III data. These are branches in which the reproducibility of raw data can be easily, though perhaps not inexpensively, tested. Reproducibility is usually used to increase the statistics, i.e., to increase the statistical credibility or quality of the data. For instance, in elementary particle physics differential cross-section values for a given process usually are the only kind of



Fig. 1. Contributing to the geophysical data explosion, Soviet space assembly Soyuz-22 before launch. This mission, in September 1976, provided massive remote sensing data obtained with photographic systems that included a Lett (Jane) MKF-6 multispectral camera covering the range 4800-8400 Å (installed in the top compartment of the spacecraft assembly shown in the figure). The remote sensing project, named "Ridge," was conducted jointly by the USSR and the German Democratic Republic as part of the Intarkosmos program.



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R.W. Boyle

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Edited by Peter J. Hood

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Citation

Since the arrival of the space age, Herbert Friedman's whole life has been dedicated to the observation and interpretation of the space environment and its behavior. Thus, in 1949, when V-2 missiles were made available to American investigators, he began his experiments at the Naval Research Laboratory by adapting laboratory instruments to measure in space the solar ultraviolet and X-ray light and its absorption in the high atmosphere. Then he went on to investigate the effect of the solar radiation on the ionosphere. His interest in space geophysics—the influence of sun on earth—has never flagged since. He has always shown superb judgment in choosing experiments which were both scientifically significant and achievable. Hence, he was led to pioneering discoveries in geophysical undersounding.

Behind Herbert Friedman's leadership of the E.O. Hubble Center for Space Research at the Naval Research Laboratory there developed a number of teams exploring sun, earth, and the interplanetary medium from space with discriminating understanding. He has encouraged them to collaborate so with outside scientific teams that now it is sometimes difficult to keep track of the myriad cooperative relationships. Moreover, in recent years he has continued to advance the cause of geophysical investigation, including especially the use of observations from space, through advocacy of support for fundamental space geophysical investigations as member and chairman of important scientific committees and commissions. As a publicist for good science and amiable expositor of space geophysics to the wider scientific and public communities, he is well known. In brief, he so well exemplifies one who has made outstanding contributions to fundamental geophysics and [one] who stands for unselfish cooperation in research that award of the William Bowie Medal to him is specially fitting. Thus he is a worthy and distinguished geophysicist who appropriately joins the lineage of previous Bowie Medal recipients renowned for their accomplishments and influence.

When, after some 9 years at the Naval Research Laboratory, Herbert Friedman turned from laboratory (X-ray) research to space experiment, the initial rocket observations were of fundamental geophysical significance. Thus he conducted the first space observations of the role of solar X rays, Schumann ultraviolet, and Lyman alpha in the production of the ionosphere. He was principal contributor to the study of the relationship between solar flux variability and ionospheric behavior over a solar cycle (1949–60). He also was responsible for the theoretical prediction and first observations of the role of solar hard X rays in producing ionospheric breakup. Next, the fundamental contribution of the first ray ultraviolet monitoring satellite—SOLRAD-1 (1960)—initiated the whole new age of space environment monitoring.

Then Friedman provided the first theoretical model of the E and F region ionosphere based on rocket observations of X rays, the extreme ultraviolet, and the dissociation of molecular oxygen in the high atmosphere. He first observed the ultraviolet airglow from rockets; the Lyman-alpha airglow of the high atmosphere was discovered; it revealed the hydrogen geocorona. He identified the Lyman-beta hy-

drogen glow of the night sky, principal input to support of the night-sky ionosphere in the E region. He provided the first X-ray photograph of the sun and thus showed the relationship between X-ray active regions and microwave radioheliograms.

In fostering unselfish cooperation in research, Herbert Friedman's contributions have been marked by knowledgeable and far-reaching vision and continuing diligence. He has been especially influential in developing international cooperative programs in solar-terrestrial research. He served as chairman of the Inter-Union Committee for Solar-Terrestrial Physics (IUCSTP) during the IUGY and was primarily instrumental in obtaining its conversion to the Special Committee for STP (SCOSTEP), which has essentially permanent status in the International Council of Scientific Unions. He served as first president of SCOSTEP, 1968–74, and initiated the organization of the International Magnetospheric Study (IMS).

In recent years, Herbert Friedman has chaired the Geophysics Research Board (GRB), the Committee on Solar-Terrestrial Relationships (CSTR), and several studies under National Academy of Sciences/National Research Council auspices which have contributed to the health and development of this field of solar-terrestrial research. Through the years he played a key role in developing the scientific cooperative missions of the Committee on Space Research (COSPAR) as a member of the executive committee, 1961–75, and as vice-president, 1971–75.

Herbert Friedman is a multitalented man (we neglect to talk of his proficiency in art and love for tennis and classical music) whose creative fundamental research and unswerving effort over a whole lifetime to foster cooperation in research are hardly adequately summarized by the outline sentences above. Nor have we referred to his service on editorial boards or to his role in publicizing and describing geophysics to a wider audience (he serves as editor and writer for the section "Review of Space Science" in the AIAA journal *Astronautics and Aerodynamics*). These activities, even if significant, are peripheral to the present comment. However, it is important to emphasize that united with his excellence and cooperative dedication in geophysical research is a personable demeanor which is forthright, understanding, and amiable, but persistent. Indeed Herb Friedman's approach to problems has always been never to give up on the important efforts but always to identify the simpler but most significant next step to take. That has led him to major geophysical research discoveries and the most valuable progress in cooperative ventures. Those are the core attributes of a Bowie medalist.

This citation was prepared by Phillip Menge and presented by Norman F. Ness.

Acceptance

William Bowie took a prominent part in shaping the destiny of the American Geophysical Union in its beginning. It was he who advocated enlarging the membership from committee size to a full-fledged scientific society, so that the original 50 members grew to our present AGU of 13,000. Deemed by his contemporaries as a man of the most inspiring presence and persuasion, he used his extraordinary talents to help create the International Union of Geodesy and Geophysics and set the course of international cooperation in geophysics for generations to follow. It is indeed an inspiration as well as a great honor to receive the Bowie Medal.

My scientific career began when William Bowie ended. In these last 40 years, science has become the main cultural phenomenon of our time. It pleased me to discover that William Bowie was a member of the astronomy section

of the National Academy of Sciences. In the grand unification of nature science today, all disciplines come together so that we have a "melting pot" sociology of scientists in which physicists, geophysicists, and astrophysicists are amalgamated.

Geologists await the physicists' determination of the lifetime of the proton to decide on the symmetry of the universe. Solar-terrestrial physicists speculate on the connection between the missing solar neutrinos and the possible influence of the sun on climate.

The study of magnetospheres is bounded by scale sizes that range from the compact pulsars to the hundred-thousand light-year dimensions of head-tail galaxies. In between are the varied personalities of the solar system magnetospheres.

So sensitive are the techniques of radio interferometry and laser ranging that they measure the tiny ellipses of continents—movements no faster than the growth of a fingernail. Incredibly, we detect signatures of micron dimensions on neutron stars thousands of light-years distant.

How baffling is the ultimate puzzle of who we are, where we come from, and why we are here! Life's origins are entwined in the processes of molecule building in interstellar space, the role of exploding stars in triggering the condensation of primordial gas clouds, and the evolution of ecologically favorable life zones in planetary environments. The hot surfaces and murky atmosphere of Venus, the turbulent clouds of Jupiter, the rings of Saturn, the dead soil of Mars, and the dying whisper of microwaves left over from the Big Bang are parts of a cosmic tapestry in which we search for answers. How fortunate that we can pool interdisciplinary talents to join this search.

William Bowie's spirit of scientific cooperation is more appropriate now than ever before.

Herbert Friedman

Meetings

Delegates to IUGG Association Meetings

U.S. scientists planning to attend the 21st General Assembly of the International Association of Seismology and Physics of the Earth's Interior (IASPEI), to be held in London, Ontario, July 30–31, 1981; the Fourth Scientific Assembly of the International Association of Geomagnetism and Aeronomy (IAGA), to be held August 3–15, 1981, in Edinburgh, Scotland; the Third Scientific Assembly of the International Association of Meteorological and Atmospheric Physics (IAMAP), to be held August 17–22, 1981, in Hamburg, Germany, should notify A. F. Spilhaus, Jr., Secretary of the U.S. National Committee for IUGG, 2000 Florida Avenue, N.W., Washington, D.C. 20009, so that they can be placed on the official list of delegates from the United States to these meetings. ☐

1982 COSPAR Meeting

The first bulletin for the 24th plenary meeting and associated activities of COSPAR contains preliminary program plans for the symposia and workshops scheduled for the meeting, May 17–June 3, 1982, in Ottawa, Ontario, Canada. Information on travel, registration, and accommodations is also included. A second bulletin, to be published in September, will contain more detailed information.

Advance registration closes April 15, 1982, but applications for the limited funding available to participants are due February 15.

All correspondence for the meeting, including requests for the meeting bulletin, should be addressed to T. W. McGrath, Executive Member, Local Organizing Committee, XXIV COSPAR, Conference Secretariat, National Research Council, Ottawa, Ontario K1A 0R8, Canada (or telephone 613/993-0312). ☐

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FUTURE AGU MEETINGS

Chapman Conferences

Spaced Variability in Hydrologic Modeling
July 21–23, 1981, Colorado State University,
Fort Collins, Colorado
Rainfall Rates
April 27–29, 1982, Illinois Union, Urbana, Illinois

1981 Midwest Meeting
September 17–18, 1981, Radisson Hotel, Minneapolis, Minnesota

1981 Pacific Northwest Meeting
September 17–18, 1981, Central Washington University, Ellensburg, Washington

Ocean Sciences/AGU/ASLO (American Society of Limnology and Oceanography) Joint Meeting
February 16–19, 1982, St. Anthony Hotel, El Tejano Hotel, Gunter Hotel, San Antonio, Texas

Fall Meetings
December 7–11, 1981, San Francisco
December 8–10, 1982, San Francisco
December 5–9, 1983, San Francisco

Spring Meetings
May 31–June 4, 1982, Philadelphia

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